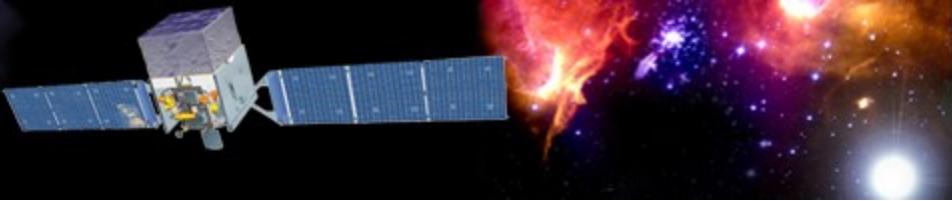




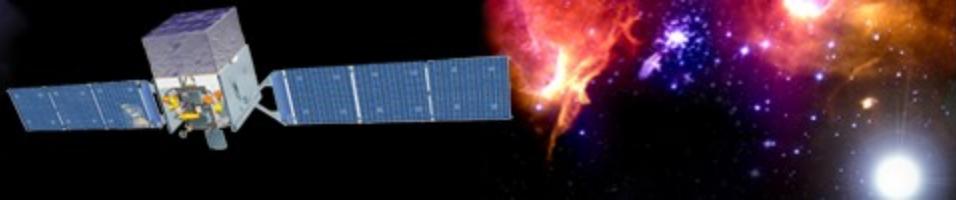
FSSC Science Tools

Source Analysis

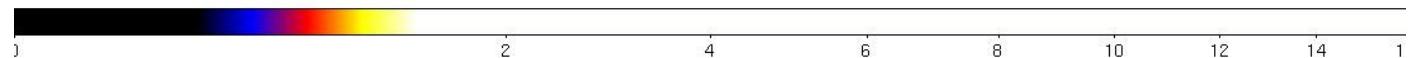
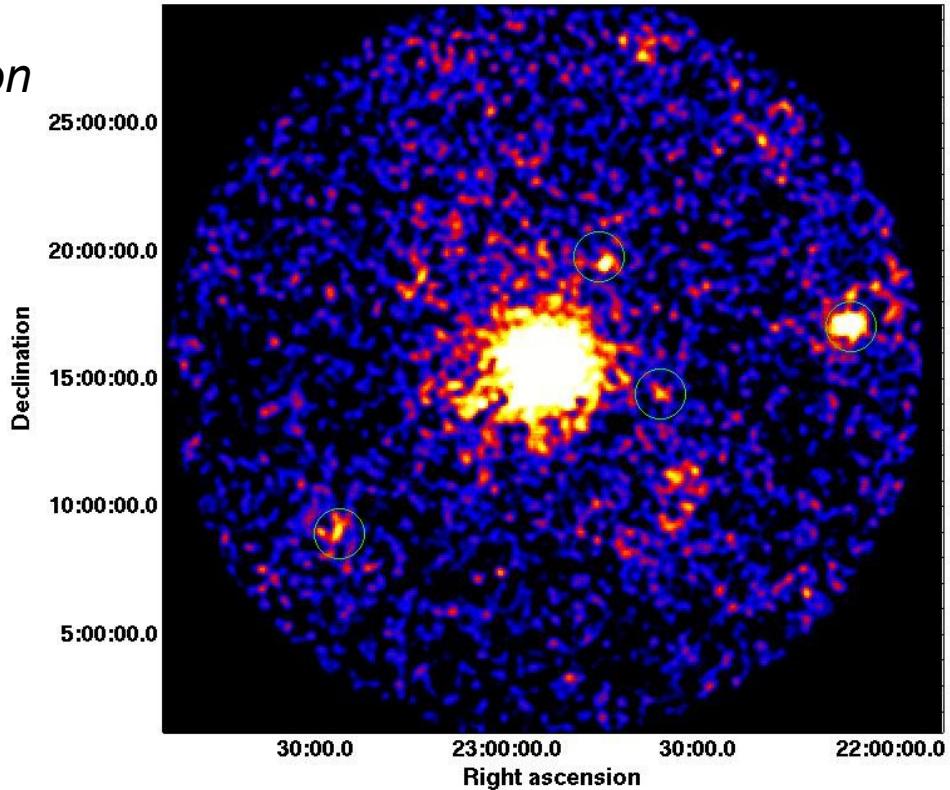


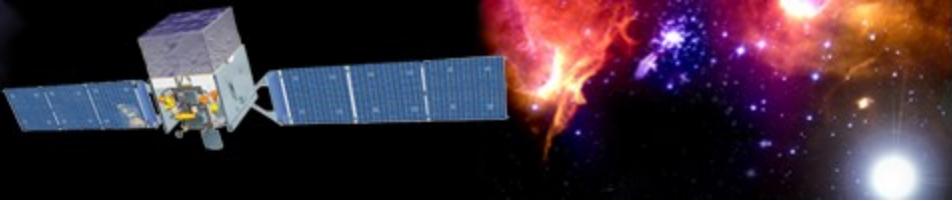
Science Tools: Documentation

- ▶ *Multi-Tier Documentation*
 - *Full set accompanies SW release*
 - *Fermi Mission Technical Handbook*
 - *Multiple levels:*
 - *Detailed analysis description ('Cicerone')*
 - *Individual tool descriptions (like fhelp)*
 - *Analysis threads (cook book examples)*



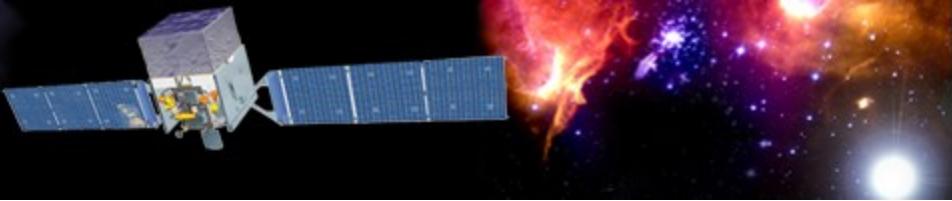
► *3c454 region*





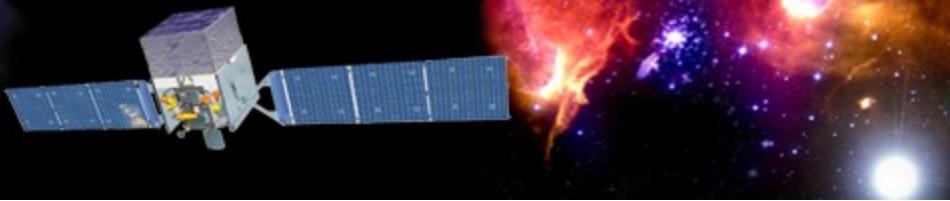
Likelihood Analysis

- ▶ *Unbinned and binned modes are available. I'll describe unbinned analysis.*
- ▶ *Several tools are needed to define the model and prepare the data*
 - *modeleditor: GUI for preparing the xml model definition file*
 - *gtselect: applies region-of-interest cuts – sky acceptance cone, energy range (0.2 – 300 GeV), time range, zenith angles (< 105°)*
 - *gtmktime: constructs good time intervals (GTIs) based on pointing information selections and zenith angle cuts*



Likelihood Analysis cont.

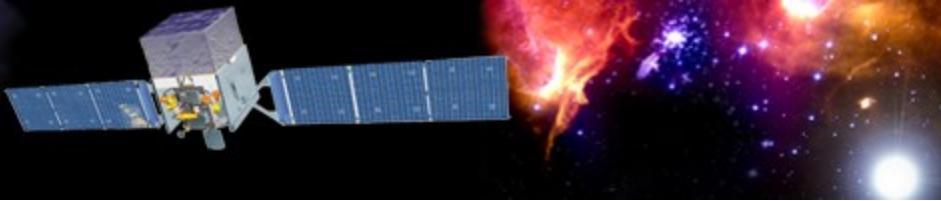
- *gtltcube*: integrates LAT livetime as a function of sky position and off-axis angle
 - *gtexpmap*: computes ROI-specific exposure maps
 - *gtlike*: fits model parameters using maximum likelihood
- Details of the method can be found in <http://fermi.gsfc.nasa.gov/ssc/data/analysis/documentation/Cicerone>



Likelihood Analysis cont.

- ## ► *Using the modeleditor GUI:*

Add sources
(point-like or diffuse)
using the drop down
menu



Likelihood Analysis cont.

Edit source name,
default f t parameters,
bounds, scaling, etc.

If a model component requires a FITS image
(e.g., Galactic diffuse,
SNR), enter the
filename here



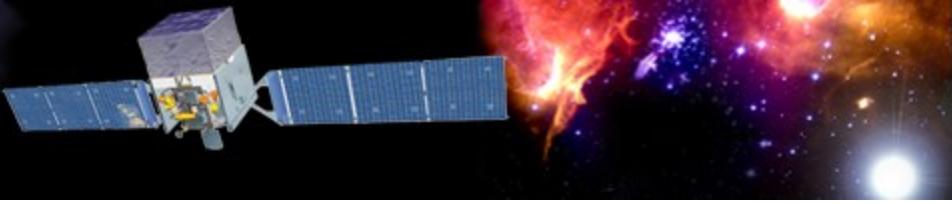
Likelihood Analysis

► *gtmktime, gtltcube*

```
>gtmktime
Spacecraft data file[3c454_SC00.fits]
Filter expression[(IN_SAA!=T) && (DATA_QUAL==1)]
Apply ROI-based zenith angle cut[yes]
Event data file[3c454_100_300000_evt01.fits]
Output event file name[3c454_100_300000_evt02.fits]
```

This removes time intervals when the ROI is intersected by the zenith angle cut

```
>gtltcube
Event data file[3c454_100_300000_evt02.fits]
Spacecraft data file[3c454_SC00.fits]
Output file[3c454_100_300000_ExpCube.fits]
Step size in cos(theta) (0.:1.) [0.025]
Pixel size (degrees)[1]
```



Likelihood Analysis

► *gtexpmap*

```
>gtexpmap
```

The exposure maps generated by this tool are meant to be used for *unbinned* likelihood analysis only.

Do not use them for binned analyses.

```
Event data file[3c454_100_300000_evt02.fits]
```

```
Spacecraft data file[3c454_SC00.fits]
```

```
Exposure hypercube file[3c454_100_300000_ExpCube.fits]
```

```
output file name[3c454_100_300000_ExpMap.fits]
```

```
Response functions[P6_V3_DIFFUSE]
```

```
Radius of the source region (in degrees)[25]
```

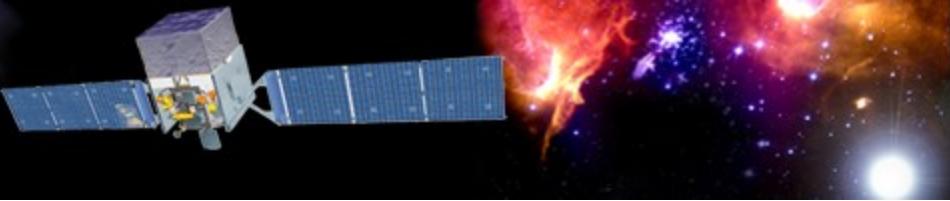
```
Number of longitude points (2:1000) [120]
```

```
Number of latitude points (2:1000) [120]
```

```
Number of energies (2:100) [20]
```

```
Computing the ExposureMap using 3c454_100_300000_ExpCube.fits
```

```
.....!
```



Likelihood Analysis cont.

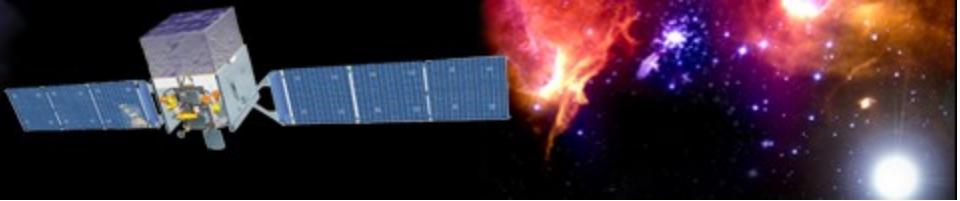
- ▶ Finally, running *gtlike*:

```
>gtlike
Statistic to use (BINNED|UNBINNED) [UNBINNED]
Spacecraft file[3c454_SC00.fits]
Event file[3c454_100_300000_evt02.fits]
Unbinned exposure map[3c454_100_300000_ExpMap.fits]
Exposure hypercube file[3c454_100_300000_ExpCube.fits]
Source model file[3c454_srcmdl.xml]
Response functions to use[P6 V3 DIFFUSE]
Optimizer (DRMNFB|NEWMINUIT|MINUIT|DRMNGB|LBFGS) [minuit]
```

...

Computing TS values for each source (6 total)
.....!

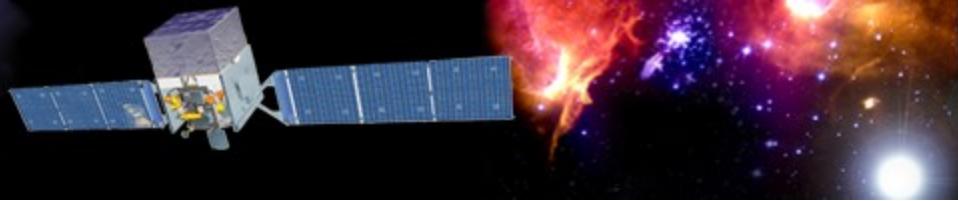
This is the xml model
file created using the
modeleditor GUI



```
GAL_v02:  
Prefactor: 1.29563 +/- 0.0426252  
Index: 0  
Scale: 100  
Npred: 16076  
  
_3c454:  
Integral: 15.6539 +/- 0.34452  
Index: 2.50803 +/- 0.0205473  
LowerLimit: 100  
UpperLimit: 300000  
Npred: 4527.59  
ROI distance: 0  
TS value: 10656.9  
WARNING: Fit may be bad in range [100, 222.696] (MeV)  
  
Total number of observed counts: 28719  
Total number of model events: 28719  
  
-log(Likelihood): 325751.9386  
  
Writing fitted model to 3c454_outmdl.xml
```

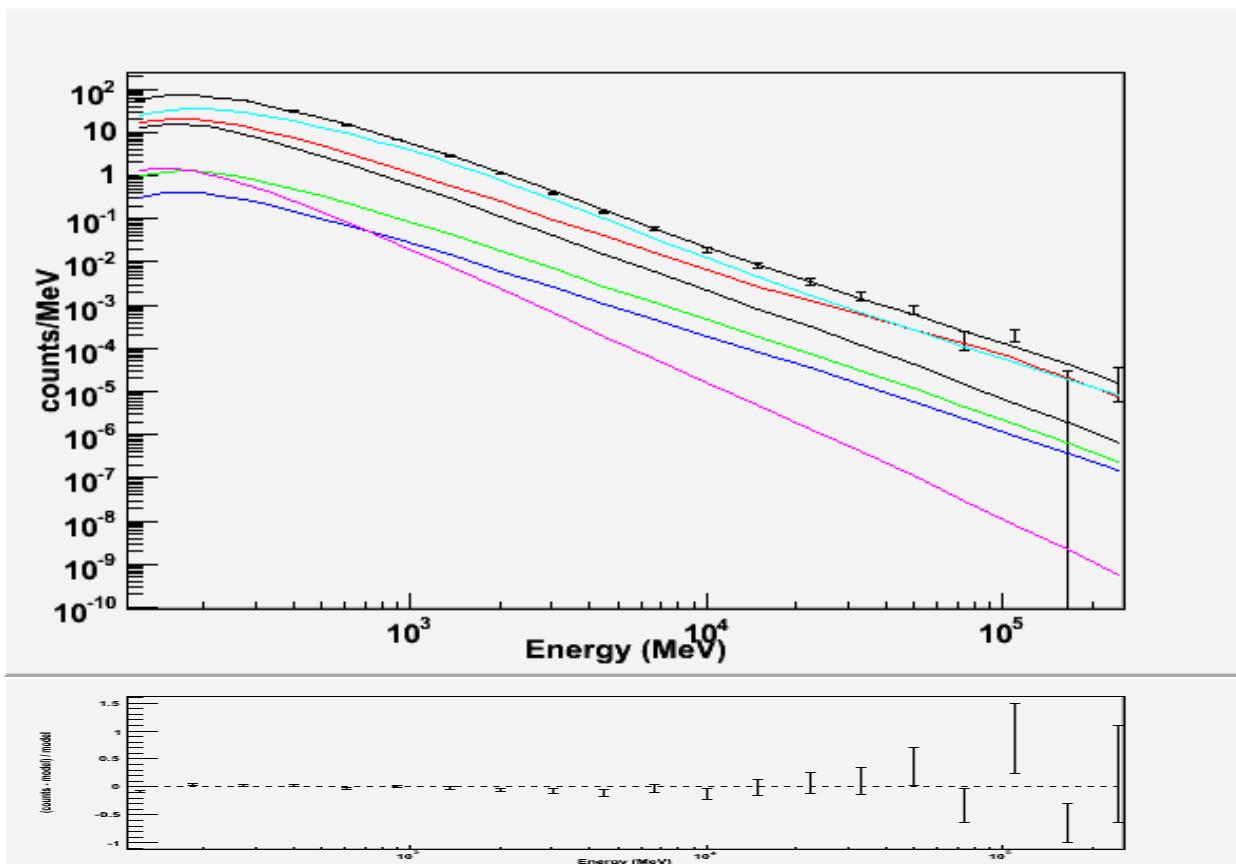
The Test Statistic (TS) is distributed as χ^2 for n dof. For a power law model TS = 25 is roughly 5σ

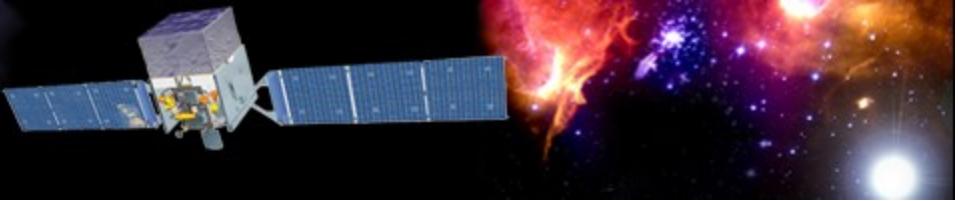
Warning messages based on Poisson probability of observed counts given the model prediction in these bands



Likelihood Analysis cont.

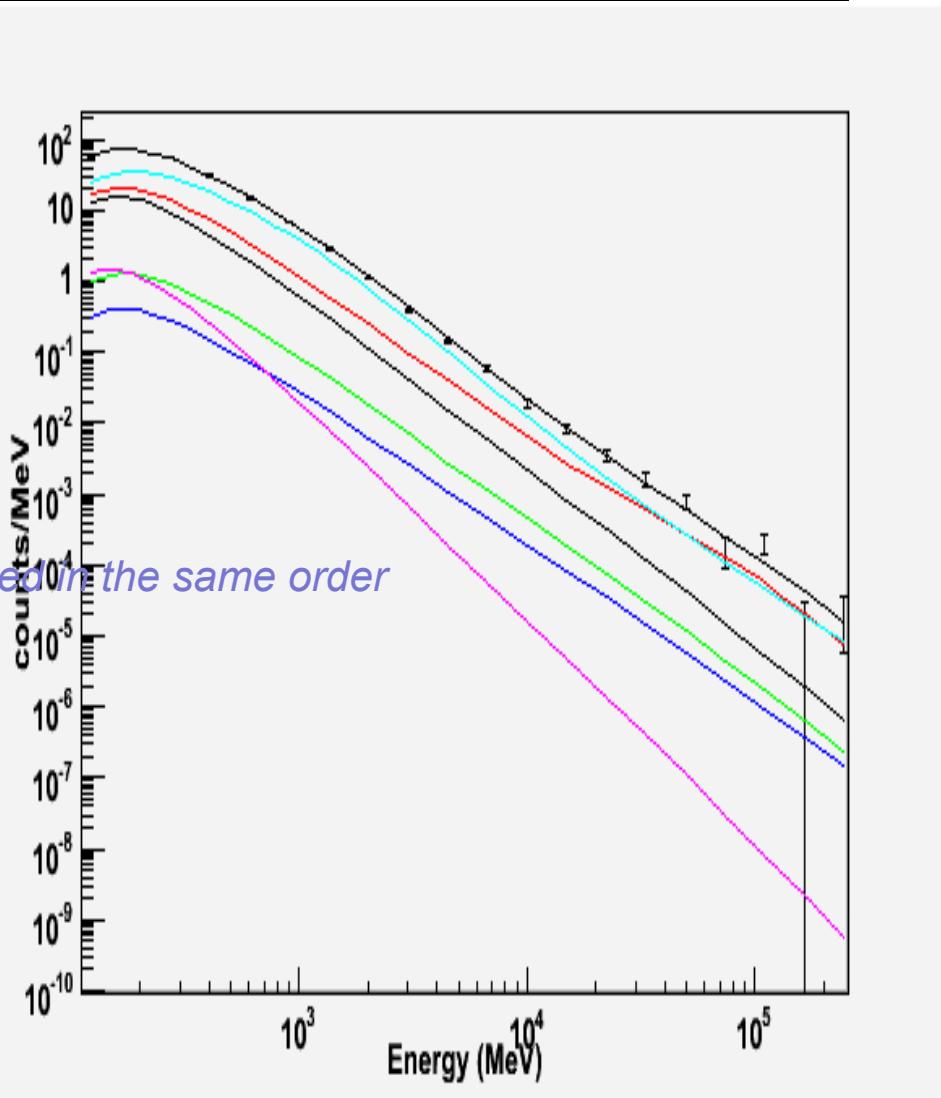
- ▶ *Plot the results (gtlike plot=yes)*

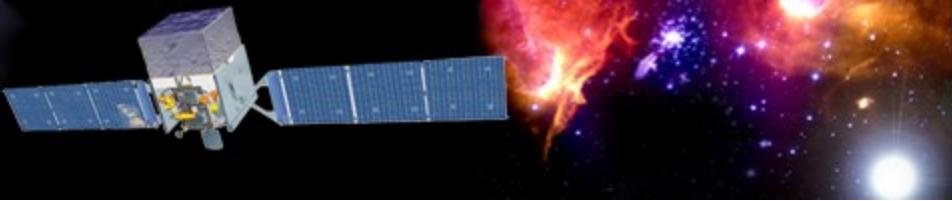




Plot Colors

- ▶ *The summed model is black*
- ▶ *The first source is red*
- ▶ *the second, green*
- ▶ *the third, blue*
- ▶ *the fourth, magenta*
- ▶ *the fifth, cyan*
 - *after this the colors are reused in the same order*

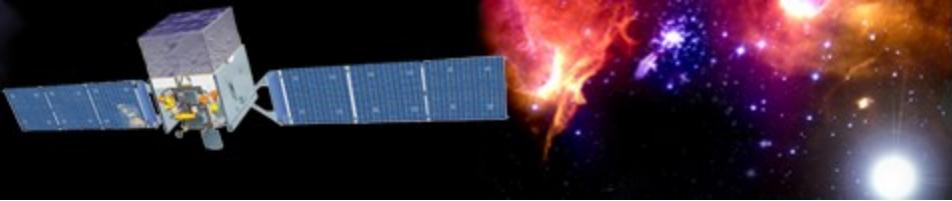




Now from Python

- ▶ *setup FSSC science tools*
- ▶ *setup python & data files*

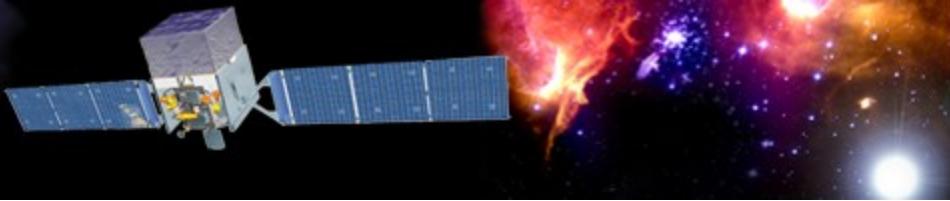
```
import glob
from UnbinnedAnalysis import *
my_obs = unbinnedAnalysis('3c454_100_300000_evt02.fits',
                           '3c454_SC00.fits',
                           expMap='3c454_100_300000_ExpMap.fits',
                           expCube='3c454_100_300000_ExpCube.fits',
                           irfs='P6_V3_DIFFUSE',
                           optimizer='MINUIT')
```



Now from Python cont.

- ▶ *setup model*

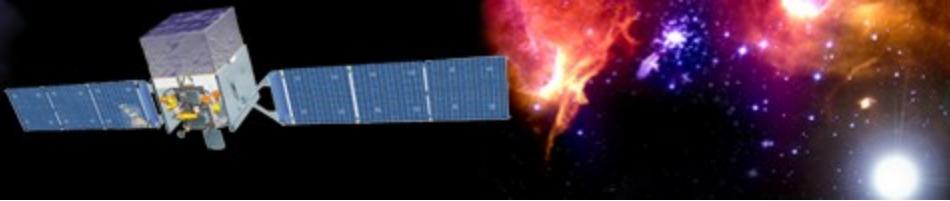
```
>>> analysis = unbinnedAnalysis(my_obs, "3c454_srcmdl.xml")
>>> print analysis
Event file(s): ('3c454_100_300000_evt02.fits',)
Spacecraft file(s): ('3c454_SC00.fits',)
Exposure map: 3c454_100_300000_ExpMap.fits
Exposure cube: 3c454_100_300000_ExpCube.fits
IRFs: P6_V3_DIFFUSE
Source model file: 3c454_srcmdl.xml
Optimizer: MINUIT
```



Now from Python cont.

- ▶ *setup model*

```
>>> analysis = unbinnedAnalysis(my_obs, "3c454_srcmdl.xml")
>>> print analysis
Event file(s): ('3c454_100_300000_evt02.fits',)
Spacecraft file(s): ('3c454_SC00.fits',)
Exposure map: 3c454_100_300000_ExpMap.fits
Exposure cube: 3c454_100_300000_ExpCube.fits
IRFs: P6_V3_DIFFUSE
Source model file: 3c454_srcmdl.xml
Optimizer: MINUIT
```



Now from Python cont.

- ▶ *fit model (verbosity=0)*

```
>>> analysis.fit(verbosity=0)
325751.93858041393
```

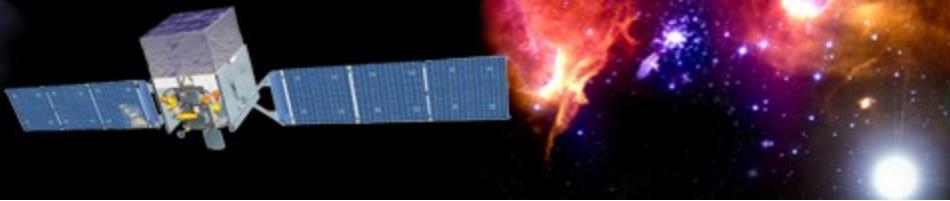
- ▶ *fit model (verbosity=1) Much more output*

```
>>> analysis.fit(verbosity=1)
*****
**      1 **SET PRINT      .000
*****
**      2 **SET NOWARN
*****

PARAMETER DEFINITIONS:
 NO.     NAME          VALUE        STEP SIZE        LIMITS
   1 'Normalizat'    .60995       1.00000      .10000E-01    10.000
   2 'Integral'      1.5776       1.00000      .10000E-03   10000.
   3 'Index'         2.3203       1.00000      1.00000      5.00000

<skip output>

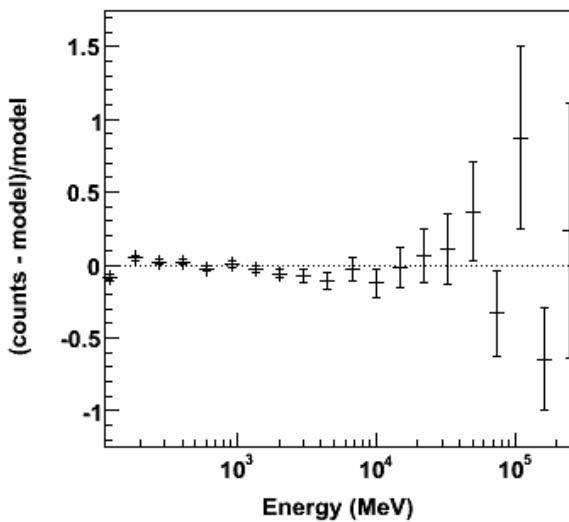
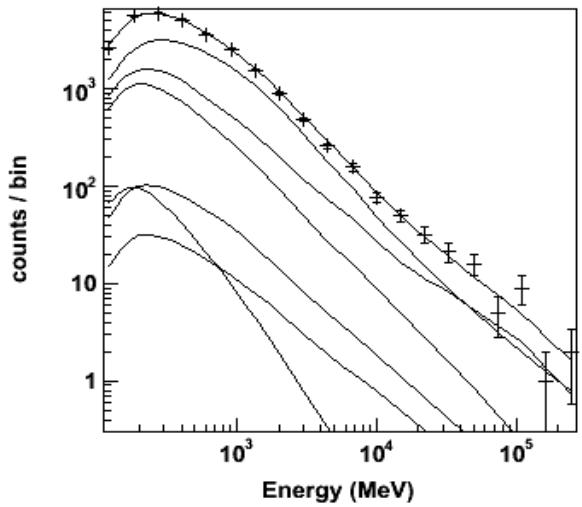
Minuit parameter uncertainties:
...
  8  0.0426301
  9  0.34453
 10  0.0205469
325751.93858041393
```

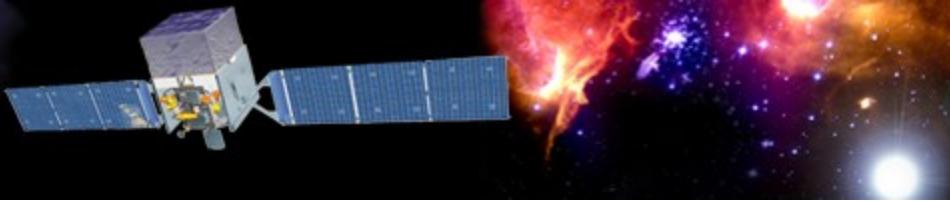


Now from Python cont.

- ▶ *Plot fitted model*

```
>>> analysis.plot()  
>>> █
```





Now from Python cont.

- ▶ *view model*

```
>>> analysis.model
EG_v02
    Spectrum: FileFunction
0      Normalization: 1.000e+00  0.000e+00  1.000e-02  1.000e+01 ( 1.000e+00)

<skip some models>

GAL_v02
    Spectrum: PowerLaw
13     Prefactor: 1.220e+00  0.000e+00  0.000e+00  1.000e+01 ( 1.000e+00)
14     Index: 0.000e+00  0.000e+00 -1.000e+00  1.000e+00 ( 1.000e+00) fixed
15     Scale: 1.000e+02  0.000e+00  5.000e+01  2.000e+02 ( 1.000e+00) fixed

_3c454
    Spectrum: PowerLaw2
16     Integral: 1.563e+01  0.000e+00  1.000e-04  1.000e+04 ( 1.000e-07)
17     Index: 2.507e+00  0.000e+00  1.000e+00  5.000e+00 (-1.000e+00)
18     LowerLimit: 1.000e+02  0.000e+00  3.000e+01  5.000e+05 ( 1.000e+00) fixed
19     UpperLimit: 3.000e+05  0.000e+00  3.000e+01  5.000e+05 ( 1.000e+00) fixed
```